

# Short Range Correlations and Spectral Functions in Exotic Nuclear Matter

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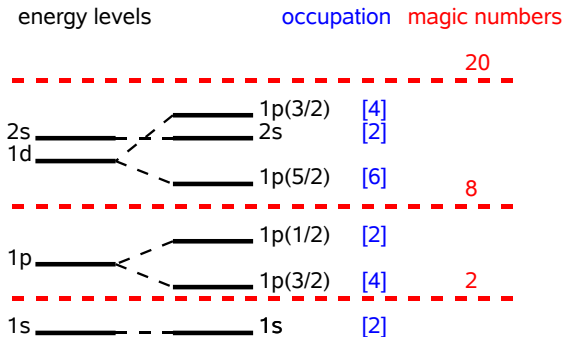
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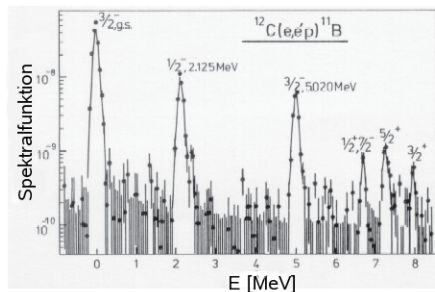
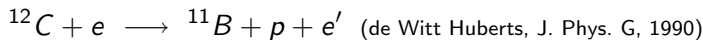
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# Shell-model

- nucleons move independently in static nuclear potential
- shell structures like in atoms



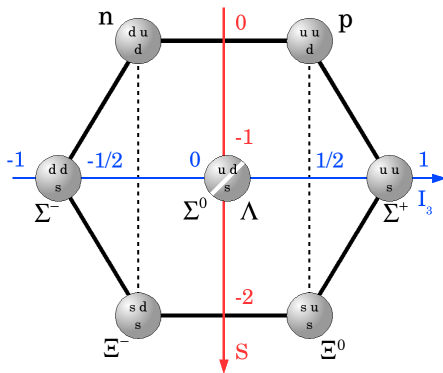
# Experimental Evidences for Correlations



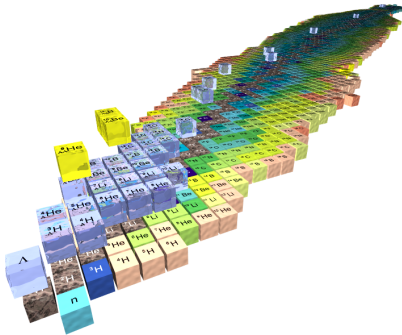
## knock-out of protons

- $E < 6$  MeV:  $1p$  shell (predictions of the shell-model)
- $E > 6$  MeV:  $2s(1/2)$  and  $1d(5/2, 3/2)$  shells (Correlations!!!)

# Exotic Matter



# Exotic Matter



(experiments for example @GSI)

## Aim

investigation of correlations in neutron rich infinite nuclear matter with strangeness admixture

⇒ better understanding of neutron stars

# One-Body Green's Function

## Definition

$$G(\vec{r}_1, t_1, \vec{r}_2, t_2) = -\frac{i}{\hbar} \left\langle T \left\{ \Psi_H(\vec{r}_1, t_1) \Psi_H^\dagger(\vec{r}_2, t_2) \right\} \right\rangle$$

propagation of an extra particle (hole) added to our system

## Temperature $T = 0$

$$\langle A \rangle = \langle \Psi_0 | A | \Psi_0 \rangle$$

## Temperature $T > 0$

thermal average over the grand canonical ensemble

$$\langle A \rangle = \frac{\text{Tr} [A e^{-\beta(H-\mu N)}]}{\text{Tr} [e^{-\beta(H-\mu N)}]}$$

## Time ordering operator

$$T \left\{ \Psi(\vec{r}_1, t_1) \Psi^\dagger(\vec{r}_2, t_2) \right\} = \begin{cases} \Psi(\vec{r}_1, t_1) \Psi^\dagger(\vec{r}_2, t_2) & \text{for } t_1 > t_2 \\ -\Psi^\dagger(\vec{r}_2, t_2) \Psi(\vec{r}_1, t_1) & \text{for } t_1 < t_2 \end{cases}$$

$$G(\vec{r}_1, t_1, \vec{r}_2, t_2) = \theta(t_1 - t_2) G^>(\vec{r}_1, t_1, \vec{r}_2, t_2)$$

$$+ \theta(t_2 - t_1) G^<(\vec{r}_1, t_1, \vec{r}_2, t_2)$$

## Particle Correlation Function

$$G^>(\vec{r}_1, t_1, \vec{r}_2, t_2) = -\frac{i}{\hbar} \langle \Psi_H(\vec{r}_1, t_1) \Psi_H^\dagger(\vec{r}_2, t_2) \rangle$$

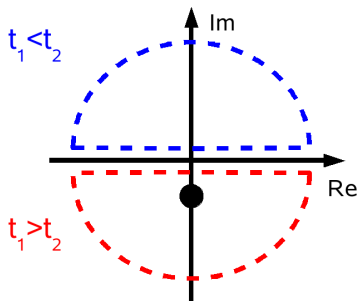
## Hole Correlation Function

$$G^<(\vec{r}_1, t_1, \vec{r}_2, t_2) = \frac{i}{\hbar} \langle \Psi_H^\dagger(\vec{r}_2, t_2) \Psi_H(\vec{r}_1, t_1) \rangle$$



## Spectral Representation of the Step Function

$$\theta(t_1 - t_2) = - \int_{-\infty}^{\infty} \frac{d\omega}{2\pi i} \frac{e^{-i\omega(t_1 - t_2)}}{\omega + i\eta}$$

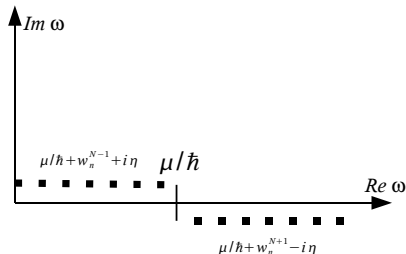


## Lehmann Representation

$$G(\vec{k}, \omega) = \frac{1}{\hbar} \sum_n \frac{\langle \Psi_0 | \Psi(\vec{k}) | \Psi_n^{N+1} \rangle \langle \Psi_n^{N+1} | \Psi^\dagger(\vec{k}) | \Psi_0 \rangle}{\omega - \mu/\hbar - w_n^{N+1} + i\eta}$$

$$+ \frac{1}{\hbar} \sum_m \frac{\langle \Psi_0 | \Psi^\dagger(\vec{k}) | \Psi_m^{N-1} \rangle \langle \Psi_m^{N-1} | \Psi(\vec{k}) | \Psi_0 \rangle}{\omega - \mu/\hbar - w_m^{N-1} - i\eta}$$

$$w_n^{N\pm 1} = \pm \frac{1}{\hbar} (E_n^{N\pm 1} - E) \quad \text{excitation energy}$$



## Thermodynamic Limit

$$G(\vec{k}, \omega) = \int_0^\infty \frac{d\omega'}{2\pi} \left[ \frac{\mathcal{A}_p(\vec{k}, \omega')}{\omega - \mu/\hbar - \omega' + i\eta} + \frac{\mathcal{A}_h(\vec{k}, \omega')}{\omega - \mu/\hbar + \omega' - i\eta} \right]$$

for sharp eigen-states:

Hole Spectral Function  $\mathcal{A}_h(\vec{k}, \omega) = 0$  for  $\omega > 0$

$$\mathcal{A}_h(\vec{k}, \omega) = 2\pi \sum_n \left| \langle \Psi_n^{N-1} | \Psi(\vec{k}) | \Psi_0 \rangle \right|^2 \delta(\omega - w_n^{N-1})$$

Particle Spectral Function  $\mathcal{A}_p(\vec{k}, \omega) = 0$  for  $\omega < 0$

$$\mathcal{A}_p(\vec{k}, \omega) = 2\pi \sum_n \left| \langle \Psi_n^{N+1} | \Psi^\dagger(\vec{k}) | \Psi_0 \rangle \right|^2 \delta(\omega - w_n^{N+1})$$

## Retarded Green's Function

$$G^{(r)}(\vec{r}_1, t_1, \vec{r}_2, t_2) = -\frac{i}{\hbar} \langle \Psi_0 | [\Psi_H(\vec{r}_1, t_1) \Psi_H^\dagger(\vec{r}_2, t_2)]_+ | \Psi_0 \rangle \theta(t_1 - t_2)$$

analytic in the upper complex plane

## Advanced Green's Function

$$G^{(a)}(\vec{r}_1, t_1, \vec{r}_2, t_2) = \frac{i}{\hbar} \langle \Psi_0 | [\Psi_H(\vec{r}_1, t_1) \Psi_H^\dagger(\vec{r}_2, t_2)]_+ | \Psi_0 \rangle \theta(t_2 - t_1)$$

analytic in the lower complex plane

## Lehmann Representation

$$G^{(r,a)}(\vec{k}, \omega) = \int_0^\infty \frac{d\omega'}{2\pi} \left[ \frac{\mathcal{A}_p(\vec{k}, \omega')}{\omega - \mu/\hbar - \omega' \pm i\eta} + \frac{\mathcal{A}_h(\vec{k}, \omega')}{\omega - \mu/\hbar + \omega' \pm i\eta} \right]$$

## Dirac Formula

$$\frac{1}{\omega \pm i\eta} = \mathcal{P} \frac{1}{\omega} \mp i\pi\delta(\omega)$$

$$G^{(r,a)}(\vec{k}, \omega) = \mathcal{P} \int_0^\infty \frac{d\omega'}{2\pi} \frac{\mathcal{A}_p(\vec{k}, \omega')}{\omega - \mu/\hbar - \omega'} + \mathcal{P} \int_0^\infty \frac{d\omega'}{2\pi} \frac{\mathcal{A}_h(\vec{k}, \omega')}{\omega - \mu/\hbar + \omega'}$$

$$\mp \frac{i}{2} \mathcal{A}_p(\vec{k}, \omega - \mu/\hbar) \mp \frac{i}{2} \mathcal{A}_h(\vec{k}, \mu/\hbar - \omega)$$

## Imaginary Part of the Retarded Green's Function

$$\text{Im} \left[ G^{(r)}(\vec{k}, \omega) \right] = \frac{1}{2} [\mathcal{A}_p(\vec{k}, \omega - \mu/\hbar) + \mathcal{A}_h(\vec{k}, \mu/\hbar - \omega)]$$

$$\equiv \frac{1}{2} \mathcal{A}(\vec{k}, \omega)$$

# Usefull Relations

## Correlation Functions

$$G^>(\vec{k}, \omega) = -e^{\beta(\hbar\omega - \mu)} G^<(\vec{k}, \omega)$$

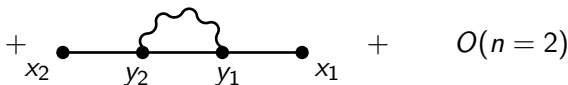
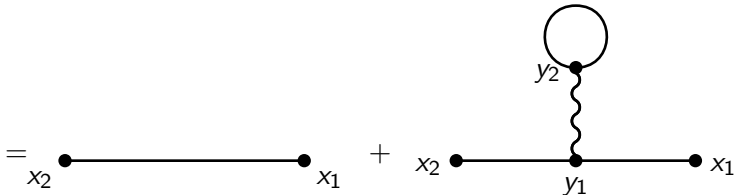
$$\begin{aligned} \mathcal{A}(\vec{k}, \omega) &= -2\text{Im} \left[ G^{(r)}(\vec{k}, \omega) \right] = i \left[ G^{(r)}(\vec{k}, \omega) - G^{(a)}(\vec{k}, \omega) \right] \\ &= i \left[ G^>(\vec{k}, \omega) - G^<(\vec{k}, \omega) \right] \end{aligned}$$

Spectral Function is a generalised density of states!!!

$$G^<(\vec{k}, \omega) = i\mathcal{A}(\vec{k}, \omega)f(\omega) \quad G^>(\vec{k}, \omega) = -i\mathcal{A}(\vec{k}, \omega)[1 - f(\omega)]$$

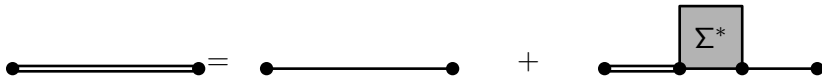
# Perturbation Expansion in the Interaction

$$G(x_1, x_2) = G_0(x_1, x_2) + G_1(x_1, x_2) + O(n=2)$$



## Dyson Equation

$$G = G_0 + G_0 \Sigma^* G$$



## Formal Solution

$$G^{(r)}(\vec{k}, \omega) = \frac{1}{\hbar\omega - \epsilon_0(\vec{k}) - \Sigma^{*(r)}(\vec{k}, \omega)}$$



## Spectral Function

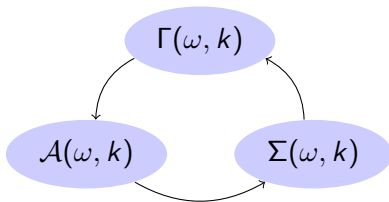
$$\mathcal{A}(k, \omega) = \frac{\Gamma(k, \omega)}{\left(\hbar\omega - \frac{\hbar^2 k^2}{2m} - \text{Re}\Sigma^{(r)}(k, \omega)\right)^2 + \left(\frac{\Gamma(k, \omega)}{2}\right)^2}$$

## Width $\Gamma > 0$

$$\Gamma(k, \omega) = -2\text{Im}\Sigma^{(r)}(k, \omega)$$

# The Model

## Self-Consistent Calculation



$$\text{Self-Energy: } \Sigma(k, \omega) = \boxed{\Sigma_{MF}(k)} + \textcircled{\Sigma_C(k, \omega)}$$

### Mean-Field Contribution



### Collisional Contribution



# Mean-Field Contribution

## Energy Density Functional

### Hartree-Fock Method

$$E = \langle \phi | (T + \frac{1}{2}V) | \phi \rangle = \int \mathcal{E}(\vec{r}) d^3r.$$

## Single Particle Energy

expansion with respect to  $p$  up to the second order

$$\begin{aligned} \epsilon_q(\rho) &= \frac{\delta \mathcal{E}}{\delta n_q(\rho)} = \frac{p^2}{2m_q} + \Sigma_q^{MF}(k, \rho) \\ &\approx \frac{p^2}{2m_q^*(\rho)} + U_q^{\text{eff}}(\rho) \end{aligned}$$

## Binding Energy

$$\frac{E}{A}(\rho) = \frac{V}{A}\rho = \frac{\mathcal{E}(\rho)}{\rho}$$

# Mean-Field Contribution

## Skyrme Parameterization

- parameterization of the N-N, N- $\Lambda$ ,  $\Lambda$ - $\Lambda$  interaction
- local interaction with 2- and 3-body-forces
- simple analytic structure
- determination of the parameters using experimental data (i.e. binding-energies)

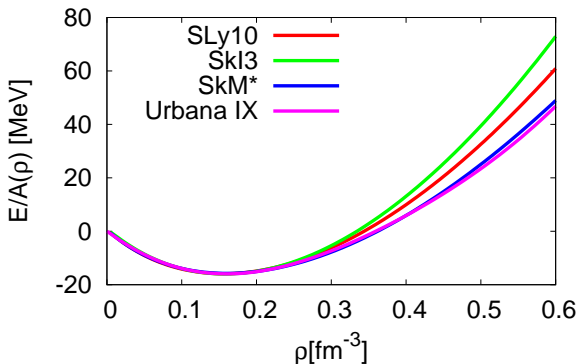
$$\begin{aligned}
 v_{ij}^{(2)}(r_1 - r_2) = & t_0(1 + x_0 P_\sigma) \delta(r_1 - r_2) \\
 & + \frac{1}{2} t_1(1 + x_1 P_\sigma) [k'^2 \delta(r_1 - r_2) + \delta(r_1 - r_2) k^2] \\
 & + t_2(1 + x_2 P_\sigma) \vec{k}' \delta(r_1 - r_2) \vec{k} \\
 & + iW_0 \vec{\sigma} \cdot [\vec{k}' \times \delta(r_1 - r_2) \vec{k}]
 \end{aligned}$$

# Equation of State

## Symmetric Nuclear Matter

- $Y = \rho_p / (\rho_n + \rho_p) = 0.5$
- saturation density  $\rho_0 \approx 0.16 \text{ fm}^{-3}$

Urbana IX: A. Akmal et al., Phys. Rev. C58, 1804-1828

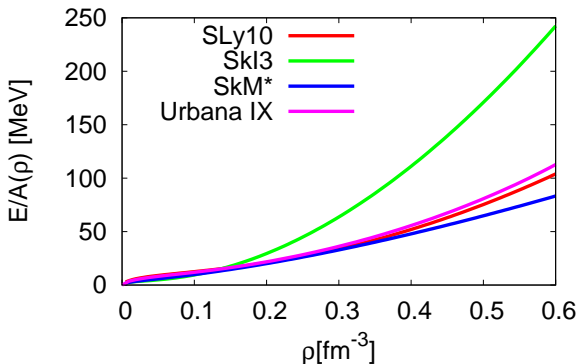


# Equation of State

## Pure Neutron Matter

- $Y = \rho_p / (\rho_n + \rho_p) = 0$
- no binding

Urbana IX: A. Akmal et al., Phys. Rev. C58, 1804-1828

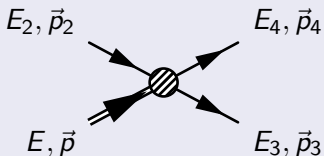


# Collisional Contribution

$$\Sigma_C(p, E) = \boxed{\Sigma^>(p, E)} + \textcircled{\Sigma^<(p, E)}$$

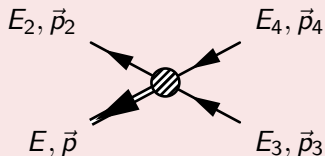
## 1-Particle-2-Hole Self-Energy

“scattering-out rate”



## 2-Particle-1-Hole Self-Energy

“scattering-in rate”



# Collisional Contribution

$$\Sigma^{\lessgtr}(p, E) \sim \int d\xi \dots \boxed{|\mathcal{M}|^2} \dots \boxed{G^{\lessgtr}(p_2, E_2) G^{\lessgtr}(p_3, E_3) G^{\lessgtr}(p_4, E_4)}$$

$$d\xi = \prod_{i=2}^4 dp_i dE_i$$

## Density of States

- particle:  $G^<(p, E) = i\mathcal{A}(p, E)f(E)$
- hole:  
 $G^>(p, E) = i\mathcal{A}(p, E)[1 - f(E)]$
- Fermi-distribution:  $f(E)$

## Local Interaction

$$|\mathcal{M}(p, E, \rho)|^2 \approx |\mathcal{M}(\rho)|^2$$

$$\Rightarrow \Sigma^{\lessgtr}(p, E) \sim \boxed{|\mathcal{M}|^2} \int d\xi \dots \boxed{G^{\lessgtr}(p_2, E_2) G^{\lessgtr}(p_3, E_3) G^{\lessgtr}(p_4, E_4)}$$



# The Spectral Function

$$\mathcal{A}(p, E) = \frac{\Gamma(p, E)}{\left(E - \frac{p^2}{2m} - \text{Re}\Sigma(p, E)\right)^2 + \left(\frac{\Gamma(p, E)}{2}\right)^2}$$

## Width

broadening of the spectral function

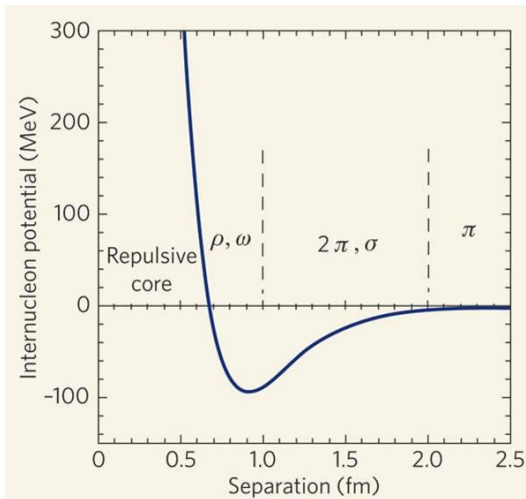
$$\Gamma(p, E) = i[\Sigma^>(p, E) - \Sigma^<(p, E)]$$

## Real Part

shifting of the peak

$$\begin{aligned} \text{Re}\Sigma(p, E) &= \Sigma_{MF}(p) \\ &+ \mathcal{P} \int \frac{dE'}{2\pi} \frac{\Gamma(p, E')}{E - E'} \end{aligned}$$

# The Nucleon-Nucleon Potential



# Landau-Migdal-Parameters

Energy Density Functional:

$$\mathcal{E}(\rho) = \sum_{qs} \int d^3p \epsilon_{qs}^0(p) n_{qs}(p) + \frac{1}{2} \sum_{qq's's'} \int d^3p d^3p' f^{qsq's'}(p, p') n_{qs}(p) n_{q's'}(p')$$

## Residual Interaction

$$f^{qsq's'}(p, p') = \frac{\delta^2 E}{\delta n_{qs}(p) \delta n_{q's'}(p')}$$

## Landau-Migdal-Parameter

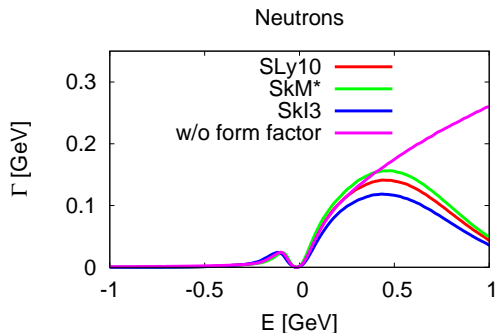
$$f_{qsq's'} = f^{qsq's'}(p_f, p'_f)$$

$p_f \approx$  Fermi-Momentum

# The Nucleon Width

## Symmetric Nuclear Matter

- $\Gamma \sim \tau^{-1}$ , direct measure for correlations
- momentum  $p = 0.100 \text{ GeV}/c$
- form factor  $F_q(E_2, E_3) = e^{(E_2 - E_3 - \epsilon_{Fq})^4 / \Lambda^4}$

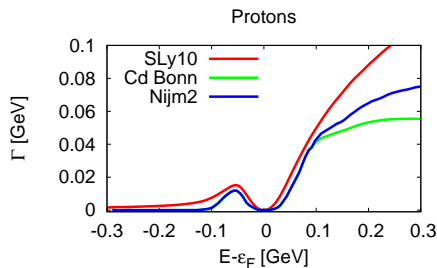
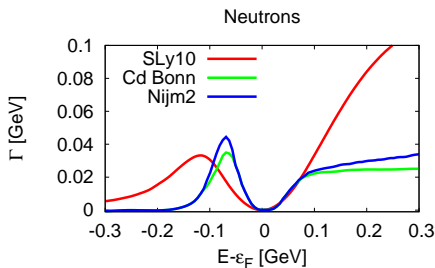


# The Nucleon Width

## Asymmetric Nuclear Matter

- asymmetry  $Y=0.25$
- density  $\rho = 0.17 \text{ fm}^{-3}$
- momentum  $p = 0.25 p_{Fq}$

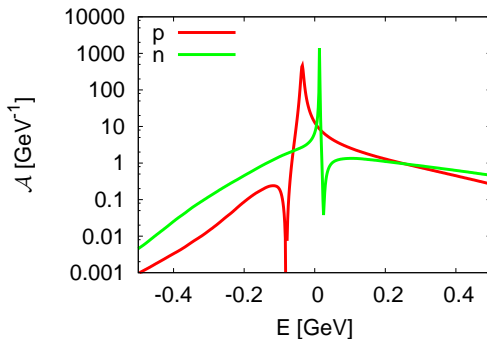
Kh.S.A.Hassaneen and H.Müther, Phys. Rev. C70 (2004)



# The Nucleon Spectral Function

## Asymmetric Nuclear Matter

- neutron rich nuclear matter  $Y = \rho_p / (\rho_n + \rho_p) = 0.1$
- $p = 0.300 \text{ GeV}/c$
- quasi-particle peak is on-shell  $\epsilon_{on}^q = \frac{p^2}{2m_q} + \Sigma_{MF}^q$



# Momentum Distribution and Spectroscopic Factor

## Momentum Distribution

$$n_q(p) = \int_{-\infty}^{\epsilon_{Fq}} \frac{dE}{2\pi} \mathcal{A}_q(E, p)$$

expansion of the real part at the on-shell point  $\epsilon_{on}$ :

$$(E - \epsilon_{on} - \text{Re}\Sigma(p, E)) \simeq (E - \epsilon_{on}) \left( 1 - \left. \frac{\partial \text{Re}\Sigma(p, E)}{\partial E} \right|_{E=\epsilon_{on}} \right)$$

## Spectroscopic Factor

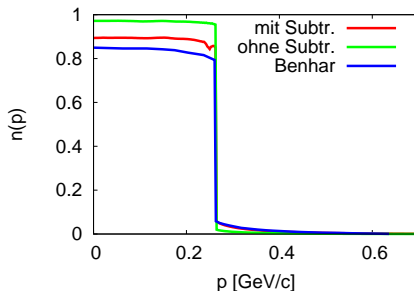
$$Z_F(p) = \left( 1 - \left. \frac{\partial \text{Re}\Sigma(p, E)}{\partial E} \right|_{E=\epsilon_{on}} \right)^{-1}$$

measure of how many particles remain in the “Quasi-particle” state

# Momentum Distribution

## Symmetric Nuclear Matter

- strength shifted above the Fermi-edge
- subtraction of long-range part important

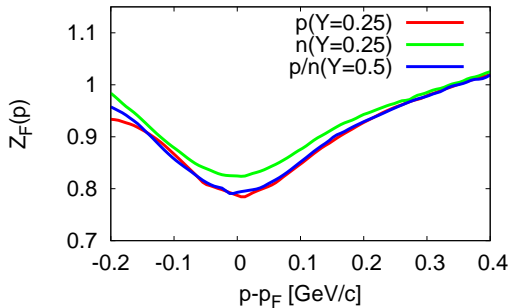




# Spectroscopic Factor

## Asymmetric Nuclear Matter

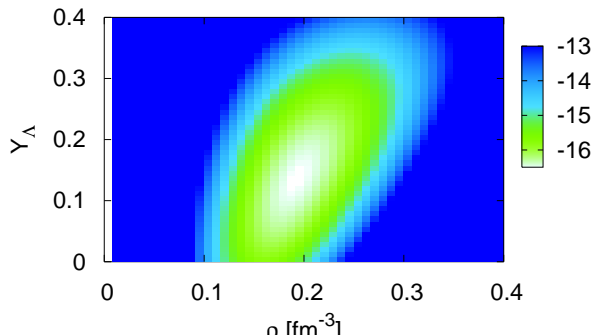
- most correlations take place close to the Fermi-edge
- less correlations for higher particle number (Pauli-blocking)



# Equation of State

## Hypernuclear Matter

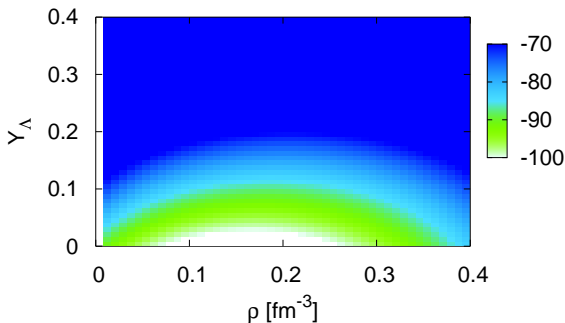
- here: equal number of proton and neutrons
- maximal binding at  $Y_\Lambda = \rho_\Lambda/\rho \approx 0.15$
- $E/A_{min} \approx -21$  MeV



# Equation of State

## Hypernuclear Matter

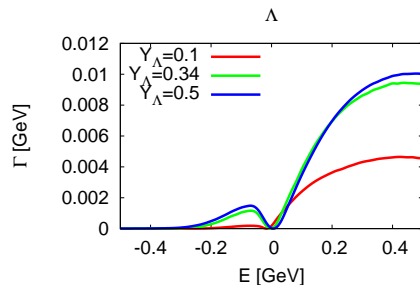
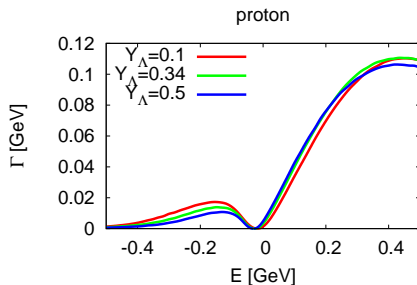
- consideration of masses:  $E/A + Y_\Lambda m_\Lambda + Y_N m_N$
- subtraction of the average mass  $m_{av} = (m_\Lambda + m_N)/2$
- maximal binding at  $Y_\Lambda = \rho_\Lambda/\rho = 0$



# Width

## Hypernuclear Matter

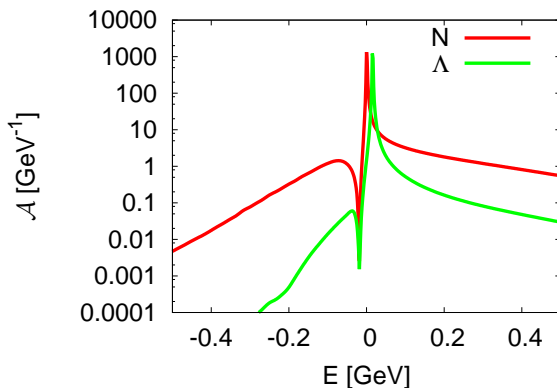
- strangeness fraction  $Y_\Lambda = 0.1$
- smaller width for  $\Lambda$ -hyperons than for nucleons



# Spectral Function

## Hypernuclear Matter

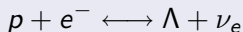
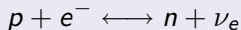
- strangeness fraction  $Y_\Lambda = 0.1$
- sharp spectral function for  $\Lambda$ -hyperons



# $\Lambda$ -Matter im $\beta$ -Equilibrium

## Condition for $\beta$ -Equilibrium

all processes are allowed, which conserve charge and Baryon number:



## Chemical Potentials

$$(\mu_e + \mu_n) - (\mu_p + m_p) = \mu_e - \mu_\nu$$

$$(\mu_n + m_n) = (\mu_\Lambda + m_\Lambda)$$

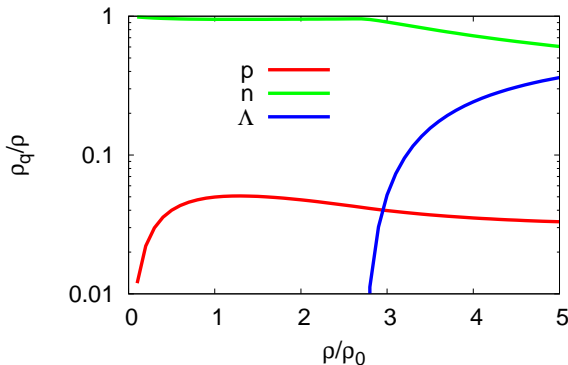
$$\text{neutrality: } n_e = n_p$$

$$\text{free Neutrinos: } \mu_\nu = 0$$

# Particle Fraction in $\beta$ -Equilibrium

## $\beta$ -equilibrium

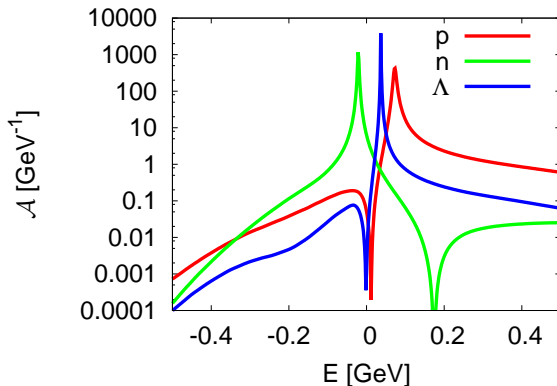
- at low densities consists  $\Lambda N$ -matter in  $\beta$ -Equilibrium nearly out of neutrons
- $\Lambda$ -formation around  $2.6\rho_0$



# Spectral Function

## $\beta$ -equilibrium

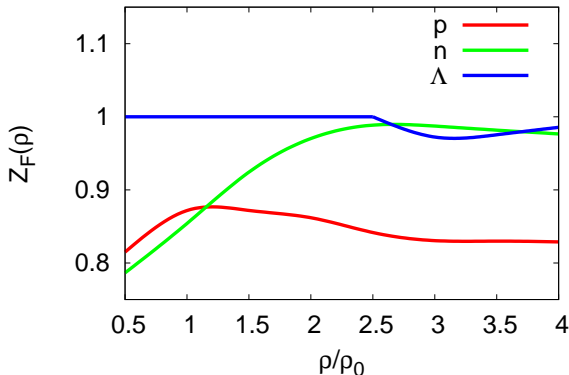
- same structure as in hypernuclear matter
- sharp spectral functions for  $\Lambda$ -hyperons and neutrons





# Spectroscopic Factor in $\Lambda N$ -Matter

- suppression of  $\Lambda$ -correlations because of  $\mathcal{M}_\Lambda < \mathcal{M}_N$
- suppression of  $n$ -correlations by Pauli-blocking



## Summary

- calculation of the spectral function in exotic nuclear matter using a self-consistent method
- energy and momentum independent interaction matrix element
- subtraction of the long-range part
  - good agreement with other many-body calculations
  - correlations do not depend on the details of the interaction
  - correlations are suppressed for high number densities

## Outlook

- calculation including  $\Sigma$ -hyperons
- finite temperature calculations